

الإسم :  
الرقم :

مادة : الفيزياء  
المدة : ساعة واحدة

**This exam is formed of three exercises in two pages.**  
**The use of non-programmable calculator is recommended.**

### **Exercise 1 (7 points)**

#### **Mechanical energy**

Consider a track ABC situated in a vertical plane as shown in document 1.

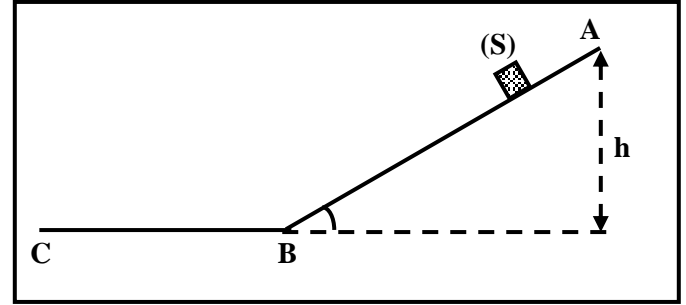
The track ABC is formed of two parts:

- an inclined part AB;
- a horizontal part BC of length  $BC = 2$  m.

A solid (S), considered as a particle of mass  $m = 0.1$  kg, is released from rest from point A.

The solid (S) is submitted to a friction force, of constant magnitude  $f$ , only along the path BC.

The horizontal plane passing through BC is taken as a reference level for gravitational potential energy.



Doc. 1

Given:

- The height of point A relative to the reference level is:  $h = 1.5$  m;
- $g = 10$  m/s<sup>2</sup>.

1) At point A:

1-1) Calculate the value of the kinetic energy  $KE_{(A)}$  of the solid (S).

1-2) Calculate the value of the gravitational potential energy  $PE_{g(A)}$  of the system [(S) - Earth].

1-3) Deduce the value of the mechanical energy  $ME_{(A)}$  of the system [(S) - Earth].

2) The solid (S) reaches point B with a speed  $V_B$ .

2-1) The mechanical energy of the system [(S) - Earth] is conserved between A and B. Why?

2-2) Deduce the value of the mechanical energy  $ME_{(B)}$  of the system [(S) - Earth] at point B.

2-3) Determine the speed  $V_B$ .

3) The solid (S) continues its motion along BC and reaches point C with a zero speed ( $V_C = 0$ ).

3-1) Calculate the mechanical energy  $ME_{(C)}$  of the system [(S) - Earth] at point C.

3-2) Calculate  $f$  knowing that  $ME_{(B)} - ME_{(C)} = f \times BC$ .

### **Exercise 2 (6.5 points)**

#### **Nuclear fusion**

If nuclear fusion were controlled in nuclear reactors, it would open prospects for sustainable economic development in the long term. Nuclear fusion usually concerns the hydrogen isotopes: deuterium  ${}^2_1\text{H}$  and tritium  ${}^3_1\text{H}$  which may merge to produce a helium nucleus  ${}^4_2\text{He}$  and a particle  ${}^A_Z\text{X}$ .

Given:

$$1u = 1.66 \times 10^{-27} \text{ kg}; \quad c = 3 \times 10^8 \text{ m/s.}$$

Nucleus or particle	${}^3_1\text{H}$	${}^2_1\text{H}$	${}^4_2\text{He}$	${}^A_Z\text{X}$
Mass (in u)	3.0160	2.0134	4.0015	1.0087

1) The nuclei  ${}^2_1\text{H}$  and  ${}^3_1\text{H}$  are isotopes. Why?

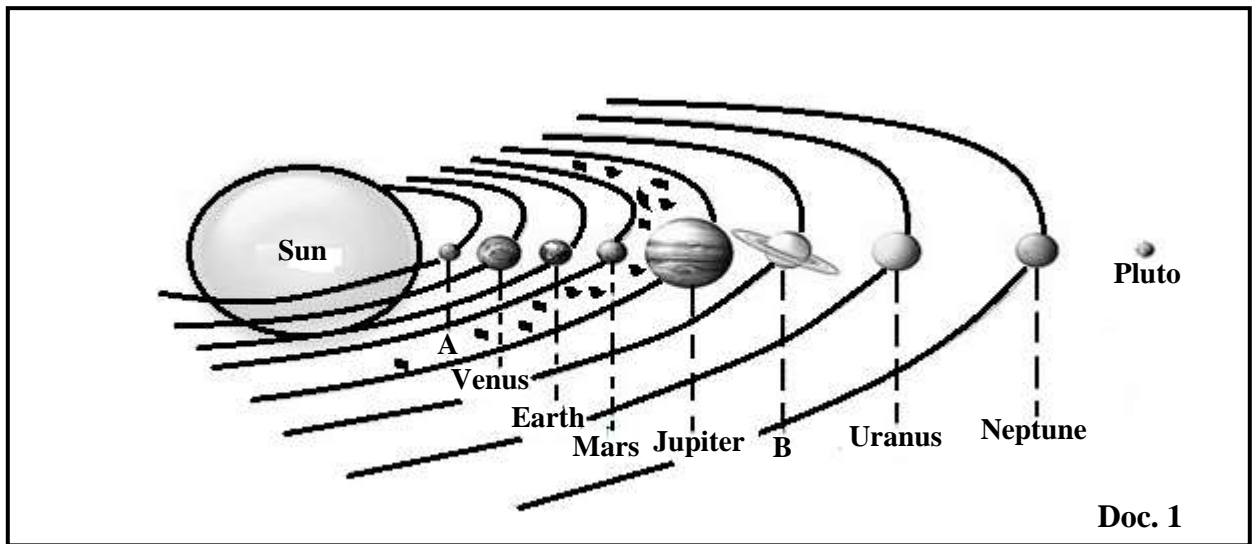
2) The fusion of  ${}^2_1\text{H}$  and  ${}^3_1\text{H}$  needs a very high temperature. Give an approximate value of this temperature.

- 3) The equation of the fusion reaction between deuterium and tritium is:  ${}^2_1\text{H} + {}^3_1\text{H} \rightarrow {}^4_2\text{He} + {}^A_Z\text{X}$ .
- 3-1) Calculate Z and A, indicating the laws used.
- 3-2) Name the emitted particle.
- 3-3) Show that the mass defect of this reaction is:  $\Delta m = 0.0192 \text{ u}$ .
- 3-4) Calculate the energy E liberated by this reaction.
- 3-5) This energy E is liberated by the fusion of one nucleus of deuterium and one nucleus of tritium of total mass of  $8.35 \times 10^{-24} \text{ g}$ . Show that the energy liberated by the fusion of 1 g of a mixture containing equal numbers of deuterium and tritium nuclei is  $E_1 = 3.4353 \times 10^{11} \text{ J}$ .
- 4) The energy liberated by the fission of 1 g of uranium-235 is  $E_2 = 8.2 \times 10^{10} \text{ J}$ . Deduce an advantage of nuclear fusion over nuclear fission.
- 5) Give another advantage of nuclear fusion over nuclear fission.

### **Exercise 3 (6.5 points)**

#### **The solar system**

Document 1 represents a simplified figure of our solar system.



- 1) The planet "A" is the closest planet to the Sun.
- 1-1) Name this planet.
- 1-2) Indicate the group of planets to which this planet belongs.
- 1-3) Indicate two common properties among the planets of this group.
- 2) The planets "B" and "Neptune" belong to the same group of planets.
- 2-1) Name the planet "B".
- 2-2) Indicate the group of planets to which these two planets belong.
- 3) The period of revolution of planet "A" around the Sun is  $T_A$  and that of planet "B" is  $T_B$ . Compare  $T_A$  and  $T_B$ . Justify by stating the convenient law.
- 4) A belt of solid objects exists between the orbits of Mars and Jupiter. Name these objects.
- 5) Document 1 shows that most of the planets orbit the Sun in almost the same plane. Name this plane.
- 6) Document 1 shows that the trajectories of the planets around the Sun are not circular.
- 6-1) Indicate the shape of the trajectories described by the planets.
- 6-2) Name the scientist who set out the law related to the shape of these trajectories.

**Exercise 1 (7 points)**
**Mechanical energy**

Question		Answer	Mark
1	1-1	$KE_{(A)} = \frac{1}{2} m V_A^2 = \frac{1}{2} \times 0.1 \times 0 = 0 \text{ J}$	0.5
	1-2	$PE_{g(A)} = m g h$ $PE_{g(A)} = 0.1 \times 10 \times 1.5 = 1.5 \text{ J}$	0.5 0.5
	1-3	$ME_{(A)} = PE_{g(A)} + KE_{(A)}$ $ME_{(A)} = 1.5 + 0 = 1.5 \text{ J}$	0.5 0.5
2	2-1	The mechanical energy is conserved between A and B since there is no friction.	0.5
	2-2	$ME_{(B)} = ME_{(A)} = 1.5 \text{ J}$	0.5
	2-3	$ME_{(B)} = PE_{g(B)} + KE_{(B)}$ $PE_{g(B)} = 0 \text{ J}$ since B is at the reference level of gravitational potential energy. $ME_{(B)} = 0 + \frac{1}{2} m V_B^2$ , so $V_B = \sqrt{\frac{2ME_{(B)}}{m}}$ , then $V_B = \sqrt{\frac{2 \times 1.5}{0.1}} = 5.5 \text{ m/s}$	0.5 0.5 0.5
3	3-1	$ME_{(C)} = PE_{g(C)} + KE_{(C)}$ $PE_{g(C)} = 0 \text{ J}$ ; since C is at the reference level of gravitational potential energy and $KE_{(C)} = 0 \text{ J}$ since $V_C = 0$ . $ME_{(C)} = 0 + 0 = 0 \text{ J}$	0.5 0.5
	3-2	$ME_{(B)} - ME_{(C)} = f \times BC$ , so $f = \frac{ME_{(B)} - ME_{(C)}}{BC}$ , then $f = \frac{1.5 - 0}{2} = 0.75 \text{ N}$	1

**Exercise 2 (6.5 points)**
**Nuclear Fusion**

Question		Answer	Grade
1		These nuclei have same charge number but different mass number.	1
2		100 million degrees	0.5
3	3.1	Conservation of mass number: $2 + 3 = 4 + A$ , then $A = 1$ Conservation of the charge number: $1 + 1 = 2 + Z$ , then $Z = 0$ (or student can say Soddy's laws)	1
	3.2	Neutron	0.5
	3.3	$\Delta m = m_{\text{before}} - m_{\text{after}}$ $\Delta m = m({}_1^2\text{H}) + m({}_1^3\text{H}) - m({}_2^4\text{He}) - m({}_0^1\text{n})$ $\Delta m = (2.0134 + 3.0160) - (4.0015 + 1.0087) = 0.0192 \text{ u}$	0.75
	3.4	$E = \Delta m c^2$ But $\Delta m = 0.0192 \times 1.66 \times 10^{-27} \text{ kg} = 3.1872 \times 10^{-29} \text{ kg}$ $E = 3.1872 \times 10^{-29} \times 9 \times 10^{16} = 2.86848 \times 10^{-12} \text{ J}$	1
	3.5	$8.35 \times 10^{-24} \text{ g} \rightarrow 2.86848 \times 10^{-12} \text{ J}$ $1 \text{ g} \rightarrow E_1$ Therefore $E_1 = 3.4353 \times 10^{11} \text{ J}$	0.75
4	$E_1 > E_2$ , then nuclear fusion yields more energy than nuclear fission	0.5	
5	Hydrogen is more abundant than uranium in nature Or: Nuclear fusion does not produce radioactive nuclei	0.5	

**Exercise 3 (6.5 points)****Solar System**

<b>Part</b>		<b>Answer</b>	<b>Mark</b>
<b>1</b>	<b>1-1</b>	A : Mercury	<b>0.5</b>
	<b>1-2</b>	Group of the inner planet	<b>0.5</b>
	<b>1-3</b>	They are solid planets They have similar dimensions (volume) They have similar mass They have almost same density (similar composition)	<b>0.5</b> <b>0.5</b>
<b>2</b>	<b>2-1</b>	B :Saturn	<b>0.5</b>
	<b>2-2</b>	Group of the outer planets	<b>0.5</b>
<b>3</b>		$T_A < T_B$ , since planet A is closer to the Sun than planet B. Kepler's third law: The period of revolution of a planet increases with the distance separating it from the Sun.	<b>0.5</b> <b>1</b>
<b>4</b>		Asteroids	<b>0.5</b>
<b>5</b>		The plane of the ecliptic	<b>0.5</b>
<b>6</b>	<b>6-1</b>	The form is elliptical	<b>0.5</b>
	<b>6-2</b>	Kepler	<b>0.5</b>